

Title: Central asian solar rotation

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How often does the Sun rotate?

He defined a fixed solar coordinate system that rotates in a sidereal frame exactly once every 25.38 days (Carrington, Observations of the Spots on the Sun, 1863, p 221, 244). The synodic rotation rate varies a little during the year because of the eccentricity of the Earth's orbit; the mean synodic value is about 27.2753 days.

What is solar rotation?

Solar rotation is the rotation of the Sun about its own axis. The Sun is not a solid body, but is composed of a gaseous plasma, and different latitudes rotate with different periods. The solar rotation period is 25.67 days at the equator and increases with increasing latitude, reaching 33.40 days at 75 degrees of latitude.

How does the Sun rotate at the equator?

On the surface, the Sun rotates slowly at the poles and quickly at the equator. This profile extends on roughly radial lines through the solar convection zone to the interior. At the tachocline the rotation abruptly changes to solid-body rotation in the solar radiation zone.

How many days is a solar rotation?

Solar rotation is taken to be 27.2753 days (see below) for the purpose of Carrington rotations. Each rotation of the Sun under this scheme is given a unique number called the Carrington Rotation Number, starting from November 9, 1853.

Here we reconstruct hydroclimatic changes over the past 160 years at Lake Karakul, Pamir Plateau (central Asia), using multiple, high-resolution (~0.8 yr) sedimentary proxies.

The future efforts in our understanding of solar rotation will be focused on the precise determination of the rotation rate of the solar core, tachocline, near-polar regions, and the upper convective boundary ...

Due to the impact of solar activity, the upper atmosphere over Eurasia forms a wave train-like structure, resulting in a tripolar temperature distribution pattern.

Internal rotation in the Sun shows differential rotation in the outer convective region and almost uniform rotation in the central radiative region. The transition between these regions is called the tachocline.

Central asian solar rotation

In this laboratory, students study the Sun's rotation by tracking sunspots in time-separated images taken with the CESAR solar telescope. By taking pictures every week, students can also participate in the ...

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Richard C. Carrington determined the solar rotation rate by watching low-latitude sunspots in the 1850s. He defined a fixed solar coordinate system that rotates in a sidereal frame exactly once every 25.38 ...

Total solar eclipses have occasionally left their footprints on human history for millennia, serving as spot references for the Earth's rotation speed and the solar cycle variations in the past.

Internal rotation in the Sun, showing differential rotation in the outer convective region and almost uniform rotation in the central radiative region. The transition between these regions is called the ...

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